

As an example, we apply repeated Richardson extrapolation to Euler's method (which is order 1) until we have an approximation of order 4. While this is comparable to RK4, we observe that the latter has less numerical issues.

Example 152. Suppose that $A(h)$ is an approximation of A^* of order 1.

- Combine $A(h)$ and $A(h/2)$ to get a new approximation $B(h)$ of order 2 (or higher).
- Combine $B(h)$ and $B(h/2)$ to get a new approximation $C(h)$ of order 3 (or higher).
- Spell out the approximation $C(h)$ in terms of the original approximation.
- By continuing the above, produce an order 4 approximation.

Solution.

- Since $A(h)$ is of order 1, we have $A(h) = A^* + Ch + O(h^2)$ as well as $A\left(\frac{h}{2}\right) = A^* + \frac{1}{2}Ch + O(h^2)$.

It follows that $\frac{2A\left(\frac{h}{2}\right) - A(h)}{2 - 1} = A^* + O(h^2)$. This is just Richardson extrapolation.

In conclusion, $B(h) = 2A\left(\frac{h}{2}\right) - A(h)$ is an approximation of A^* of order (at least) 2.

- We assume that $B(h) = A^* + Ch^2 + O(h^3)$ (for a new constant C). Hence, $B\left(\frac{h}{2}\right) = A^* + \frac{1}{4}Ch^2 + O(h^3)$.

It follows that $C(h) := \frac{4B\left(\frac{h}{2}\right) - B(h)}{4 - 1} = A^* + O(h^3)$. Again, this is just another Richardson extrapolation.

- We have:

$$\begin{aligned} C(h) &= \frac{4}{3}B\left(\frac{h}{2}\right) - \frac{1}{3}B(h) = \frac{4}{3}\left(2A\left(\frac{h}{4}\right) - A\left(\frac{h}{2}\right)\right) - \frac{1}{3}\left(2A\left(\frac{h}{2}\right) - A(h)\right) \\ &= \frac{8}{3}A\left(\frac{h}{4}\right) - 2A\left(\frac{h}{2}\right) + \frac{1}{3}A(h) \end{aligned}$$

- Since $C(h)$ is an order 3 approximation, the following is an order 4 (or higher) approximation:

$$\frac{8C\left(\frac{h}{2}\right) - C(h)}{8 - 1} = \frac{64}{21}A\left(\frac{h}{8}\right) - \frac{8}{3}A\left(\frac{h}{4}\right) + \frac{2}{3}A\left(\frac{h}{2}\right) + \frac{1}{21}A(h)$$

Comment. Instead of going via the intermediate approximation $B(h)$, we could have also proceeded as in the first part of the next example: starting with $A(h) = A^* + C_1h + C_2h^2 + O(h^3)$, we would look for a combination $\alpha A(h) + \beta A\left(\frac{h}{2}\right) + \gamma A\left(\frac{h}{4}\right)$ such that C_1 and C_2 get eliminated. This alternative approach is probably a bit more tedious to spell out but has the advantage that we can combine any three approximations.

Example 153. Python Let us use the Euler method as implemented in Example 142. We apply it to the IVP $y' = \frac{2}{5}y^2$ with $y(0) = 2$, which has the exact solution $y(x) = \frac{10}{5-4x}$ (note that this has a singularity at $x = \frac{5}{4}$) with $y(1) = 10$.

```
>>> def f(x, y):
    return 2/5*y**2
```

The following confirms that the error in Euler's method is $O(h)$.

```
>>> [euler(f, 0, 2, 1, 10**n)[-1] - 10 for n in range(7)]

[-6.4, -3.153039180044801, -0.5786318726117212, -0.06365106772819829,
-0.006430401713281242, -0.0006437015789977352, -6.437677910753337e-05]
```

If $A(h)$ is Euler (which is order 1) then $B(h) = 2A\left(\frac{h}{2}\right) - A(h)$ is an approximation of order 2:

```
>>> [2*euler(f, 0, 2, 1, 2*10**n)[-1]-euler(f, 0, 2, 1, 10**n)[-1] - 10 for n in range(7)]

[-4.8640000000000001, -1.0300367298617736, -0.03059704698782717,
-0.00036086415725833376, -3.672563401835305e-06, -3.678901450143712e-08,
-3.710773910370335e-10]
```

And we know how to further increase the order. The following is order 3:

```
>>> def A(n):
    return euler(f, 0, 2, 1, n)[-1]

>>> [8/3*A(4*10**n)-2*A(2*10**n)+1/3*A(10**n) - 10 for n in range(7)]

[-3.261606554763264, -0.23097082059404528, -0.0009539131127027645,
-1.1759506524100516e-06, -1.2013980921210532e-09, -4.835243316847482e-12,
2.644640062499093e-11]
```

Observe that the last approximation reveals numerical issues. These worsen if we push things to order 4:

```
>>> [64/21*A(8*10**n)-8/3*A(4*10**n)+2/3*A(2*10**n)-1/21*A(10**n) - 10 for n in range(7)]

[-1.7873721196484613, -0.03201840214106966, -1.63900245677695e-05,
-2.0771775410821647e-09, 1.0391687510491465e-12, 5.448086426440568e-12,
-2.8705926524708048e-11]
```

For comparison, let us also apply RK4 which is of order 4 as well:

```
>>> [runge_kutta4(f, 0, 2, 1, 10**n)[-1] - 10 for n in range(5)]

[-1.6372498407533236, -0.0029737693332965165, -3.5183410496131273e-07,
-3.524114333686157e-11, -6.394884621840902e-14]
```

Here, the numerical issue with the final approximation is less pronounced.

A glance at discretizing PDEs

One of the most important partial differential equations is the following equation which, for instance, models the steady-state temperature $u(x, y)$ of a region in two-dimensional space.

(Laplace equation, 2D)

$$u_{xx} + u_{yy} = 0$$

Notation. Here, for instance, $u_{xx} = \frac{\partial^2}{\partial x^2} u(x, y)$ is used to denote two partial derivatives with respect to x .

The Laplace equation is often written as $\Delta u = 0$, where $\Delta = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2}$ is the Laplace operator you may know from Calculus III.

The Laplacian Δu is also often written as $\Delta u = \nabla^2 u$. The operator $\nabla = (\partial/\partial x, \partial/\partial y)$ is pronounced “nabla” (Greek for a certain harp) or “del” (Persian for heart), and ∇^2 is short for the inner product $\nabla \cdot \nabla$.

Higher dimensions. Likewise, the Laplace equation in three dimensions is $\Delta u = 0$ with $\Delta = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$.

Comment. The Laplace equation is so important that its solutions have their own name: **harmonic functions**.

It is also known as the **potential equation** and is satisfied by electric/gravitational potential functions.

Recall from Calculus III (if you have taken that class) that the gradient of a scalar function $f(x, y)$ is the vector field $\mathbf{F} = \text{grad } f = \nabla f = \begin{bmatrix} f_x(x, y) \\ f_y(x, y) \end{bmatrix}$. One says that \mathbf{F} is a **gradient field** and f is a **potential function** for \mathbf{F} (for instance, \mathbf{F} could be a gravitational field with gravitational potential f).

One way to describe a unique solution to the Laplace equation within a region is by specifying the values of $u(x, y)$ along the boundary of that region.

This is particularly natural for steady-state temperatures profiles of a region R . The Laplace equation governs how temperature behaves inside the region but we need to also prescribe the temperature on the boundary.

The PDE with such a boundary condition is called a Dirichlet problem:

(Dirichlet problem)

$$u_{xx} + u_{yy} = 0 \text{ within region } R$$

$$u(x, y) = f(x, y) \text{ on boundary of } R$$

In general. A Dirichlet problem consists of a PDE, that needs to hold within a region R , and prescribed values on the boundary of that region (“Dirichlet boundary conditions”).